

Review of Sterile Neutrino Data

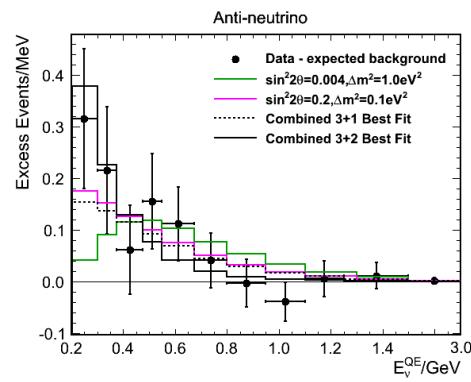
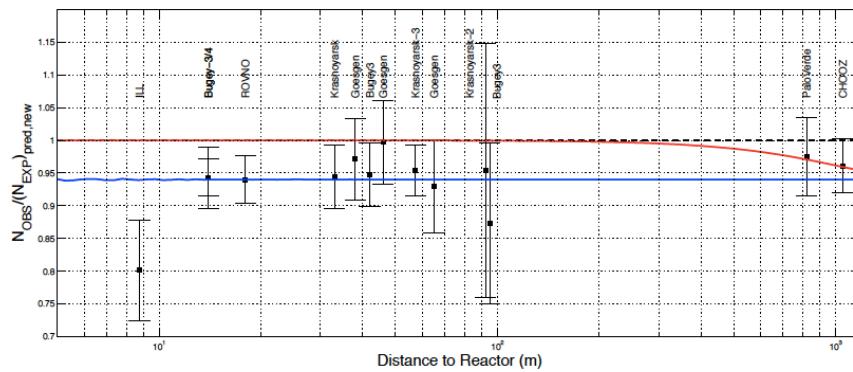
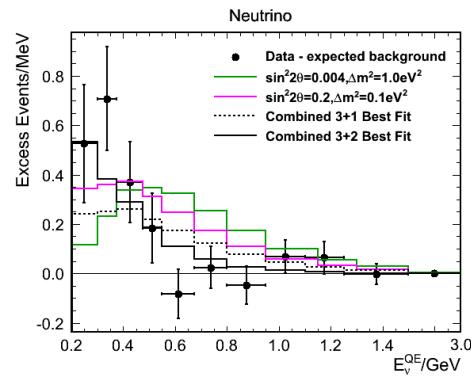
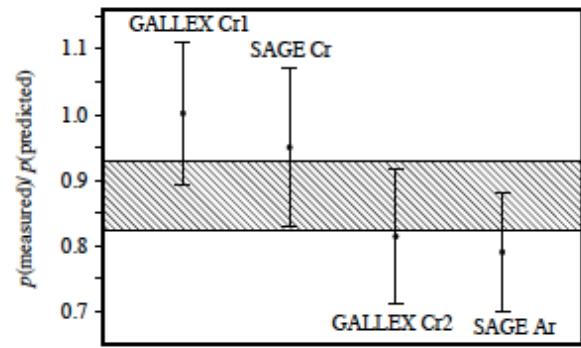
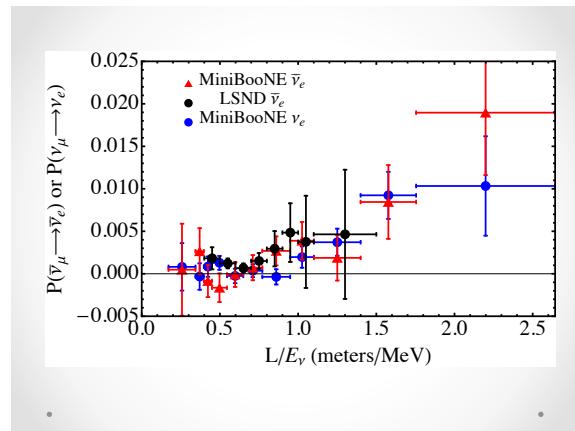
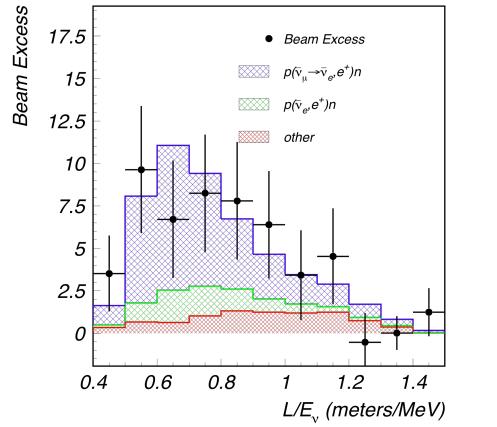
W. C. Louis

SLAC Intensity Frontier Workshop

March 6, 2013

- Short Baseline Neutrino Anomalies
- Explaining the Data with 3+N Sterile Neutrino Models
(Although there are other possibilities!)
- Global Fits to the World Data

Short-Baseline Neutrino Anomalies

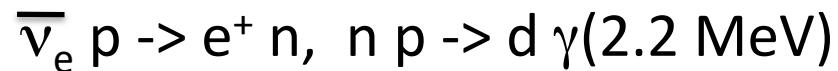
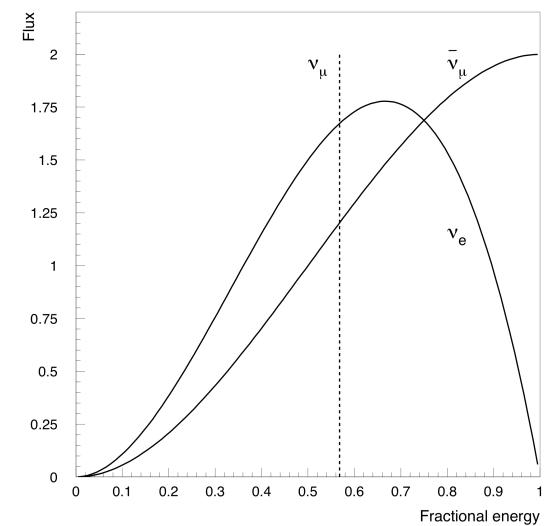
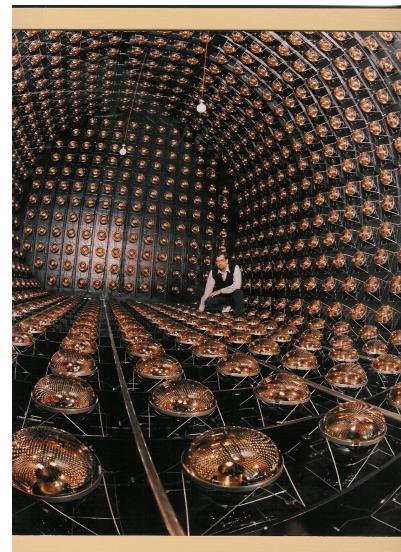
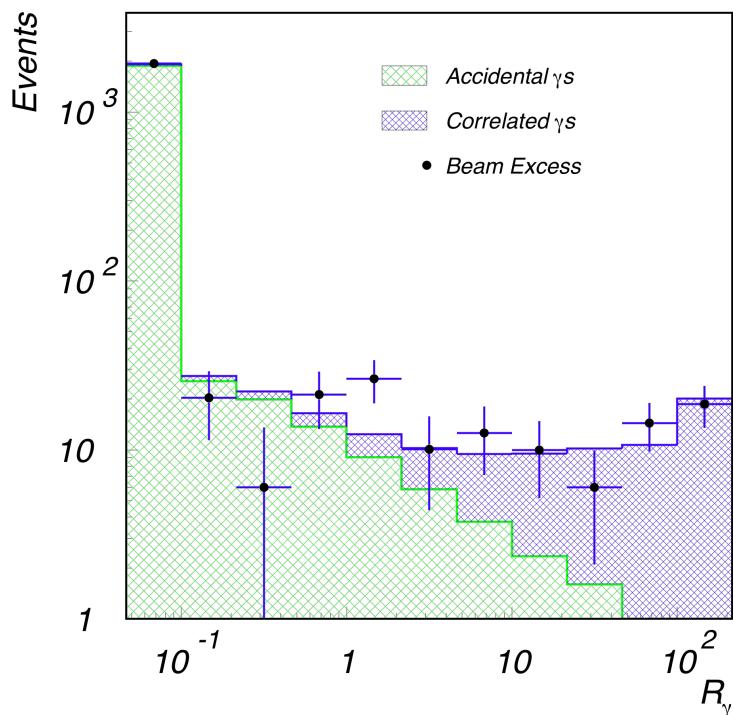


LSND Event Excess

A. Aguilar et al., Phys. Rev. D 64, 112007, (2001)

Correlated γ = 117.9+-22.4 events

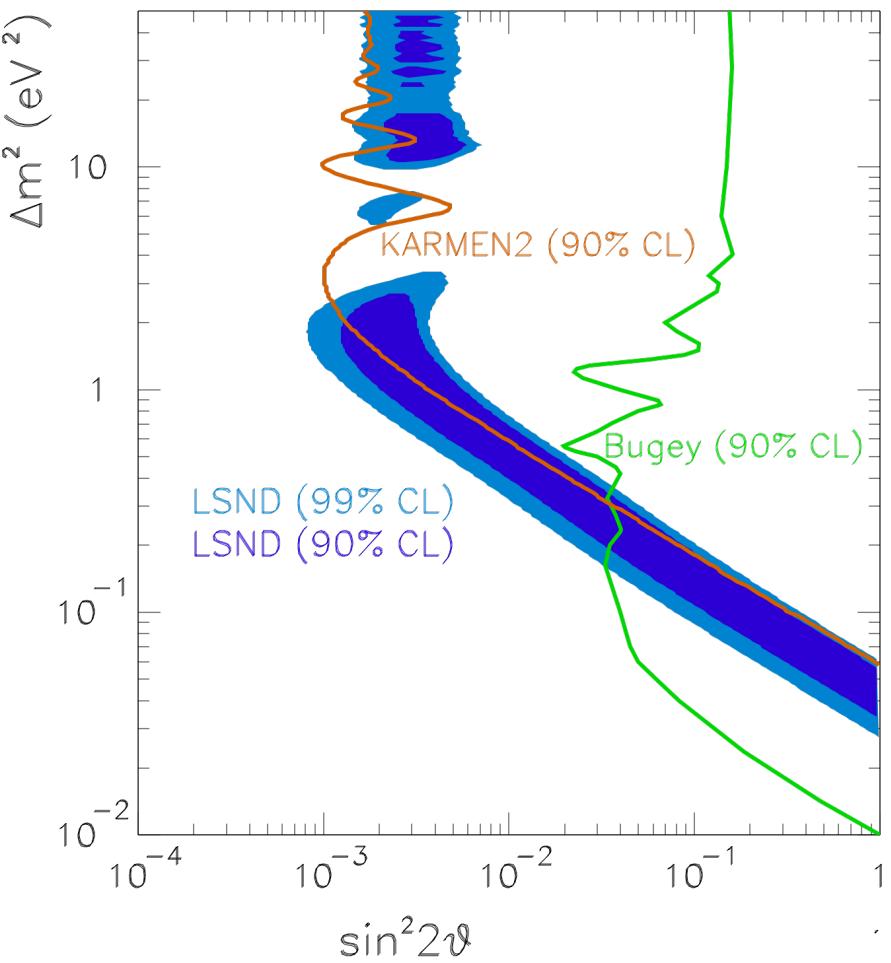
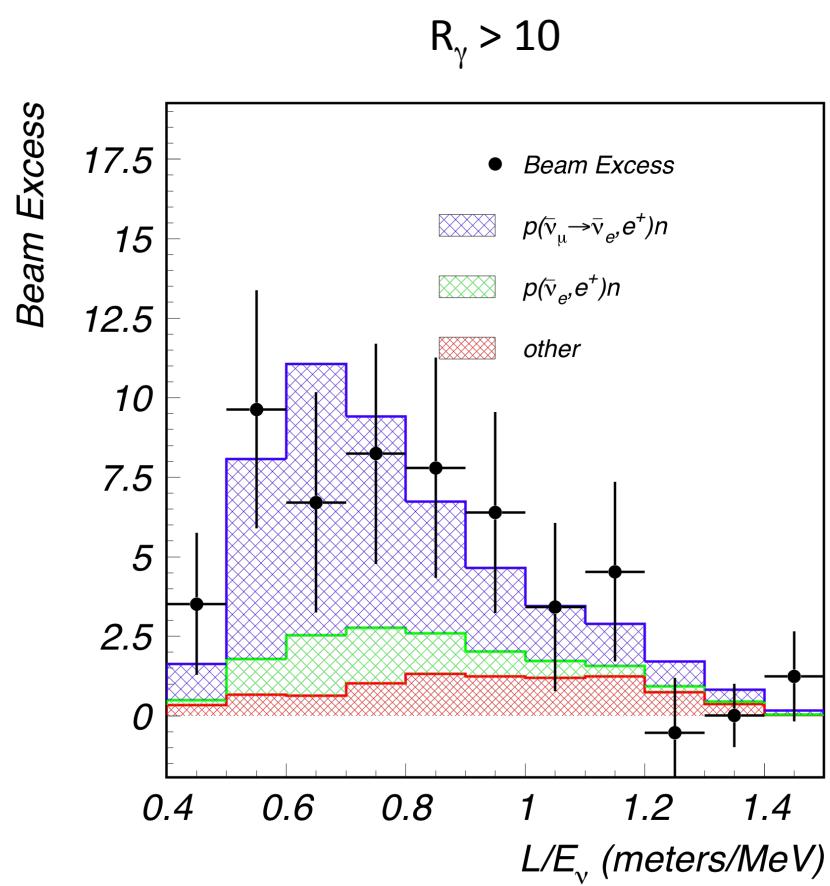
Excess = 87.9+-22.4+-6.0 events



LSND collected 28,896 C on target and observed a 3.8σ excess of events consistent with $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ oscillations, corresponding to $P_{osc} = (0.264+0.067+0.045)\%$

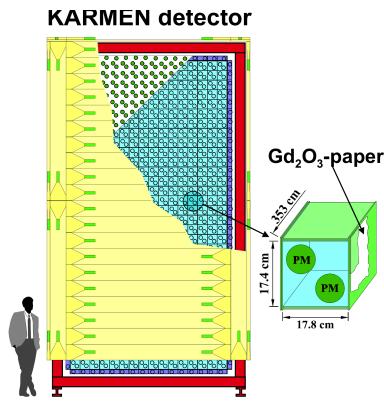
LSND Event Excesses

A. Aguilar et al., Phys. Rev. D 64, 112007, (2001)

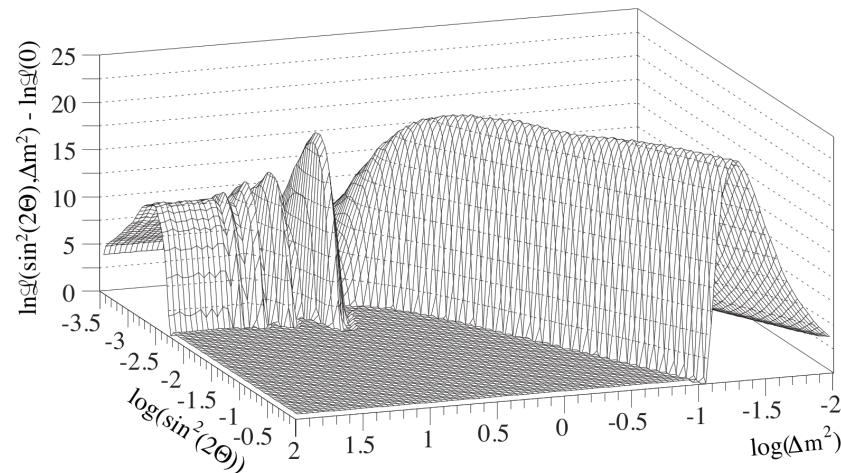


Joint LSND/KARMEN Analysis

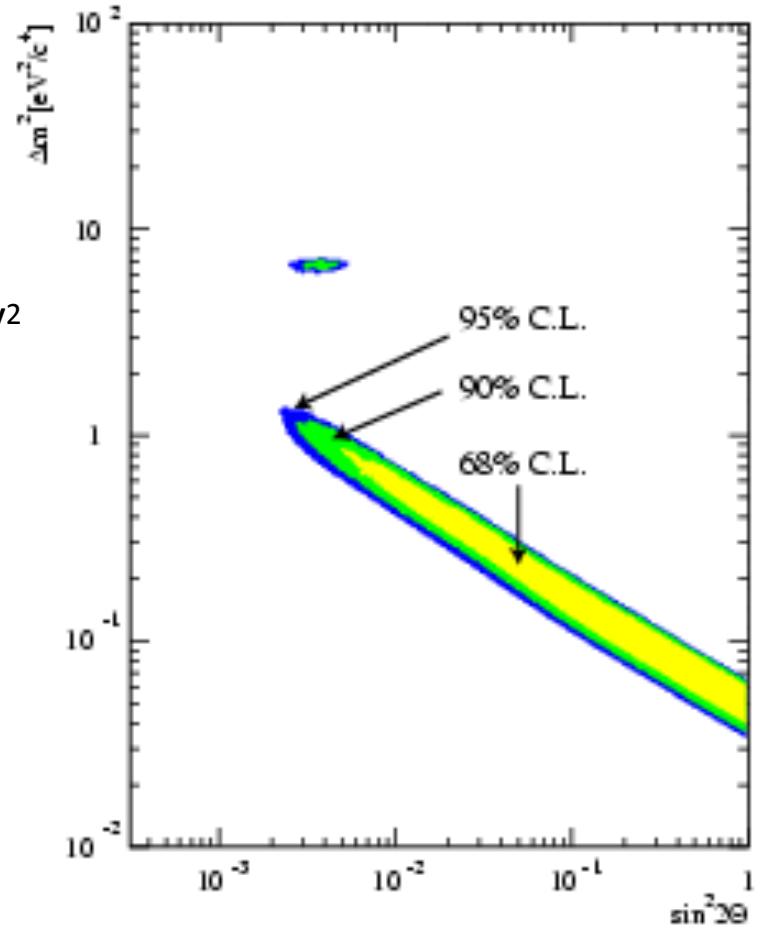
E. D. Church, K. Eitel, G. B. Mills, and M. Steidl, Phys. Rev. D66, 013001, (2002)



96% active volume of ¹²C and p
 $E = \frac{11.5\%}{\sqrt{E[\text{MeV}]}}$ $t_{\text{ISIS}} = 2\text{ns}$

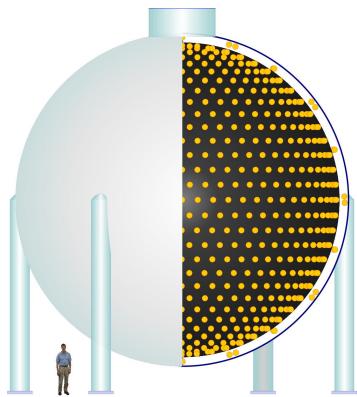


KARMEN observed no event excess; however, a joint analysis of KARMEN (17.7m) & LSND (30m) reveals a favored region of $\Delta m^2 < 1 \text{ eV}^2$

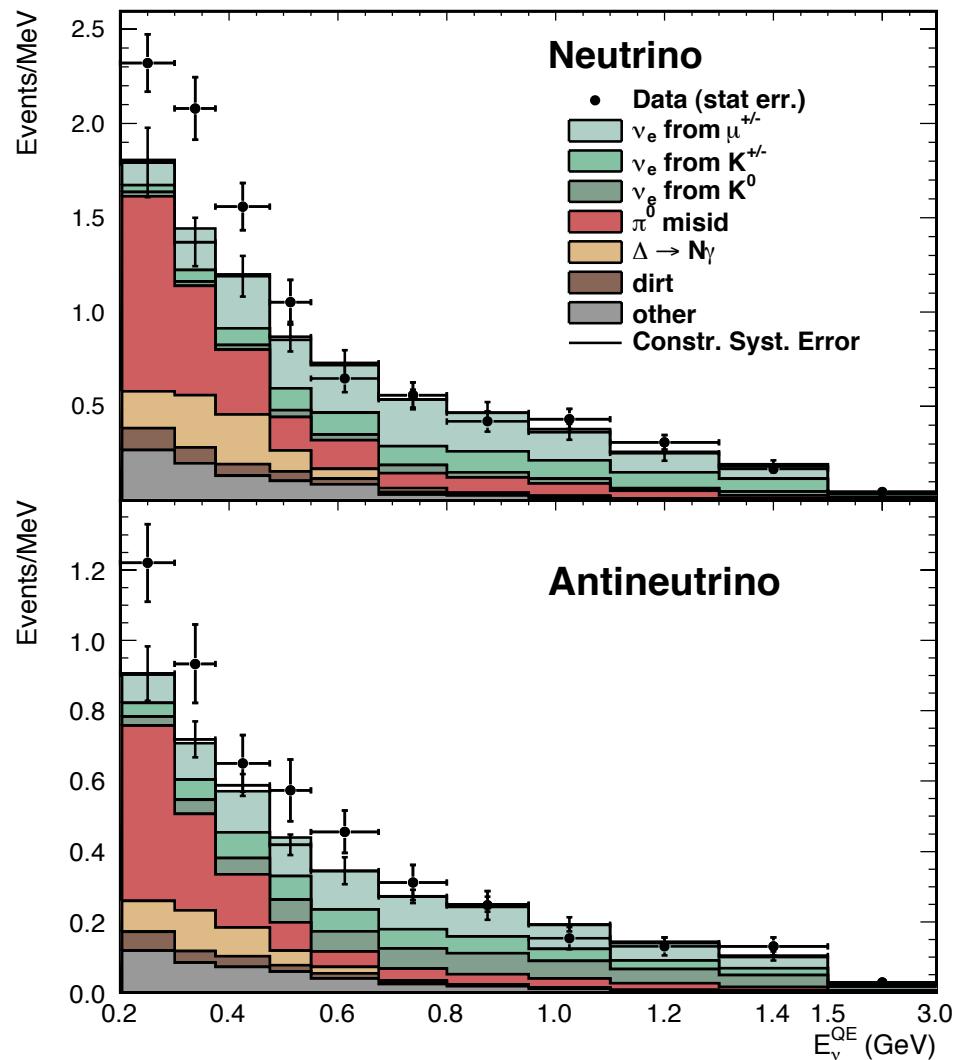


MiniBooNE ν_e Data/MC Comparison (10y)

arXiv:1207.4809

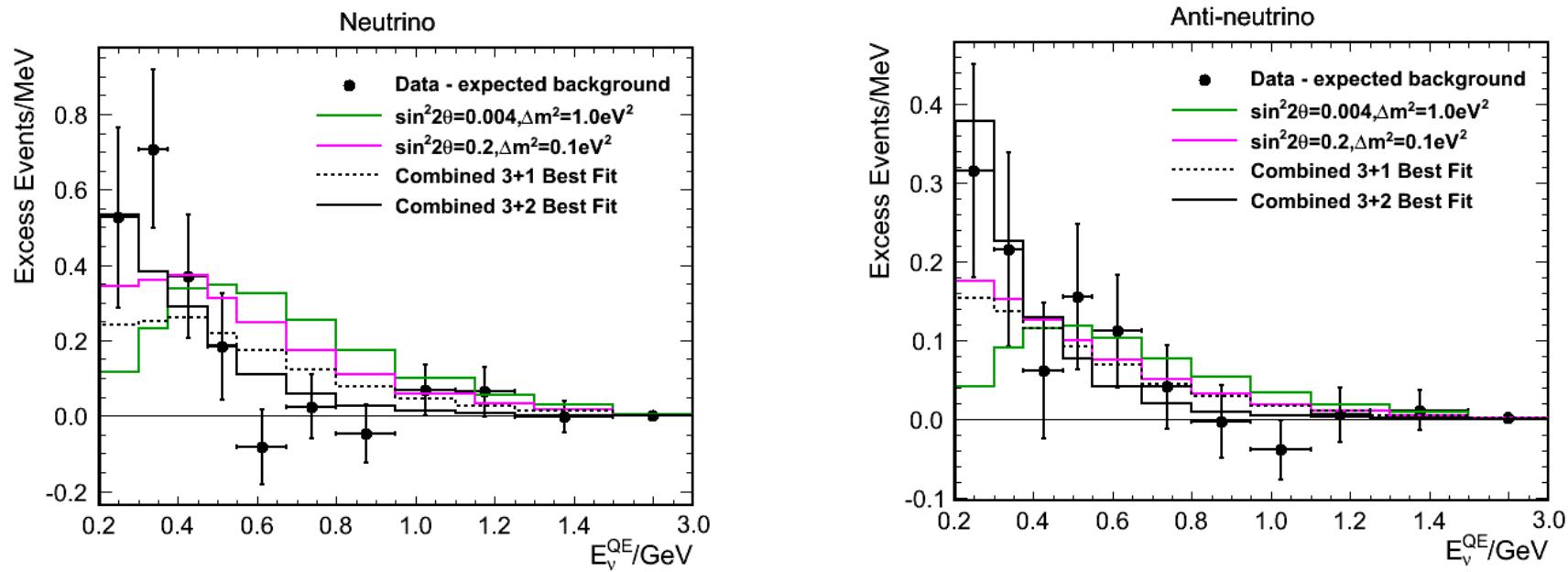


MiniBooNE observes
an excess of events
consistent with $\nu_\mu \rightarrow \nu_e$
& $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ oscillations



MiniBooNE Event Excesses

arXiv:1207.4809

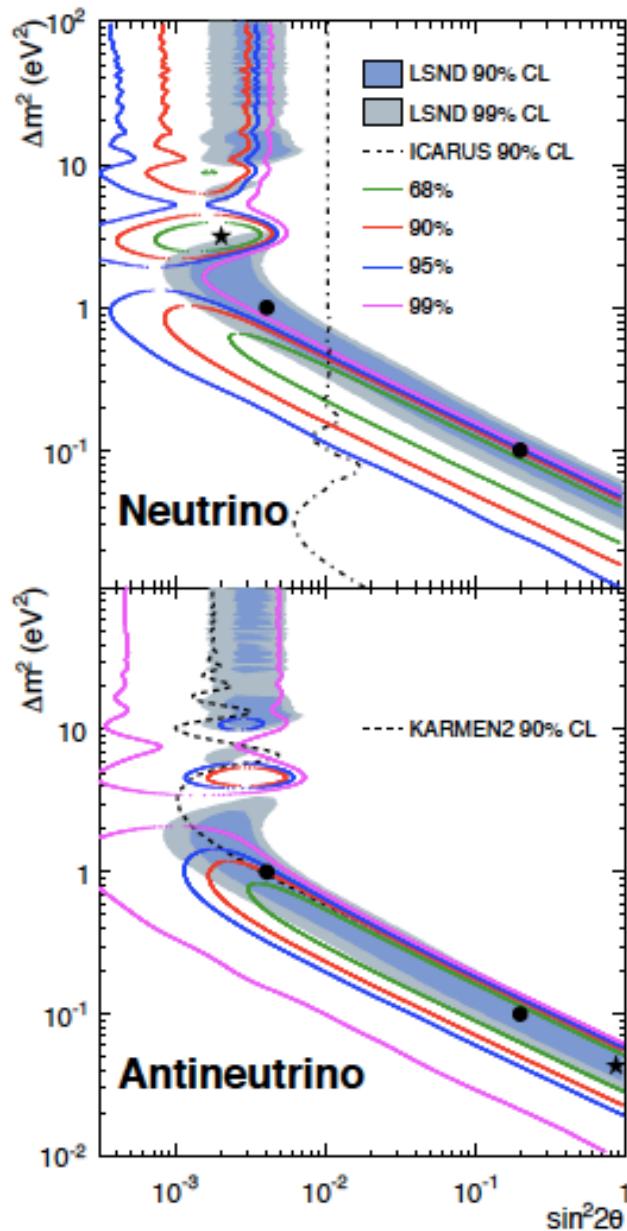


Neutrino Event Excess from 200-1250 MeV = $162.0 + -28.1 + -38.7$ (3.4σ)

Antineutrino Event Excess from 200-1250 MeV = $78.4 + -20.0 + -20.3$ (2.8σ)

Combined Event Excess from 200-1250 MeV = $240.3 + -34.5 + -52.6$ (3.8σ)

MiniBooNE Allowed Regions



Neutrino

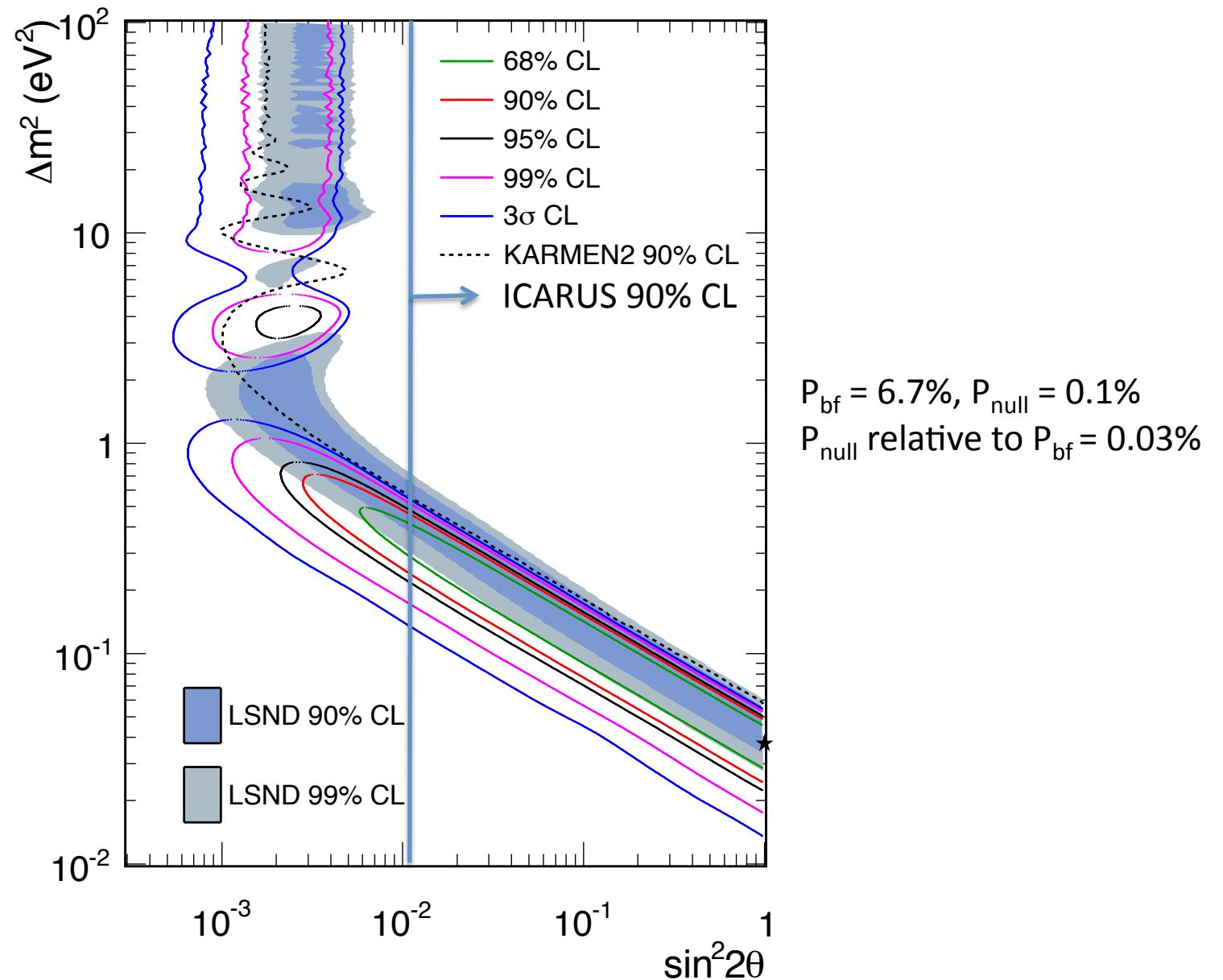
$P_{bf} = 6.1\%$, $P_{null} = 0.5\%$
 P_{null} relative to $P_{bf} = 2.0\%$

Antineutrino

$P_{bf} = 66\%$, $P_{null} = 5.4\%$
 P_{null} relative to $P_{bf} = 0.5\%$

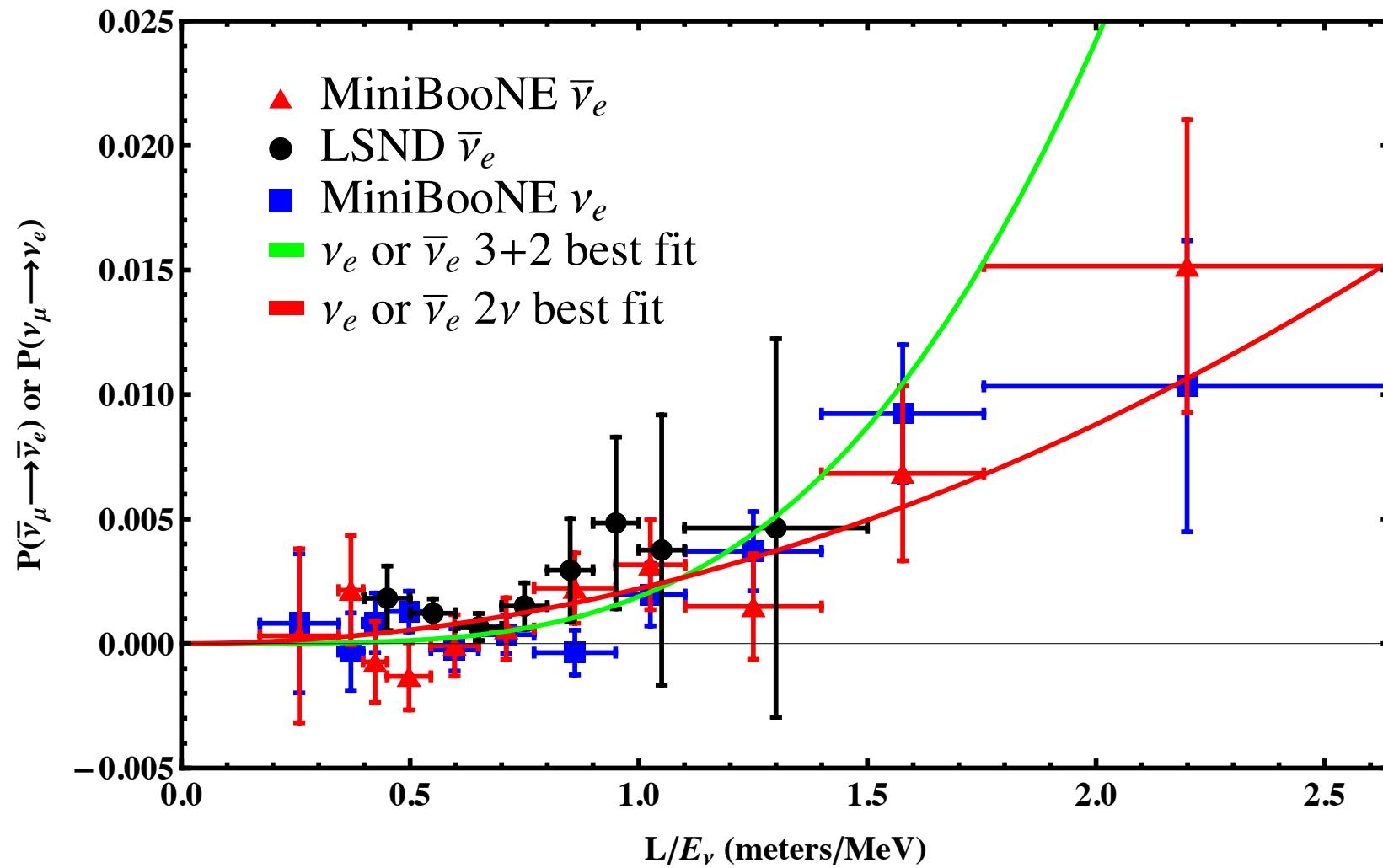
MiniBooNE Combined Neutrino + Antineutrino 2ν Fit

arXiv:1207.4809



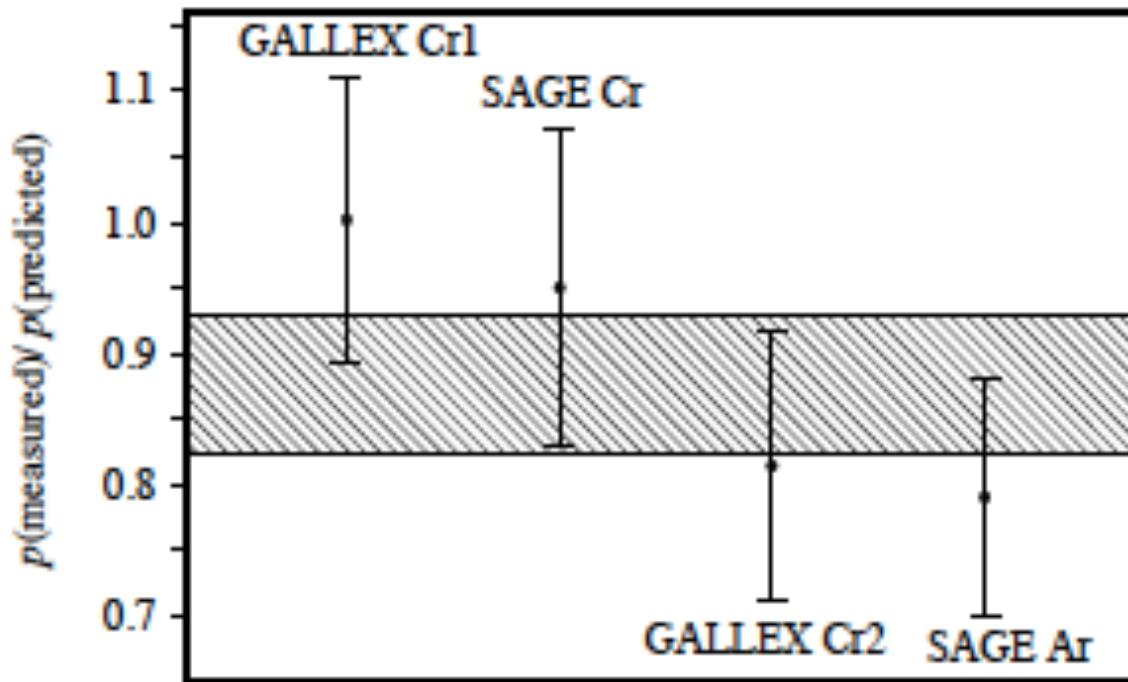
MiniBooNE L/E Distributions

arXiv:1207.4809



Radioactive Neutrino Source Anomaly

SAGE, Phys. Rev. C 73 (2006) 045805



$$R=0.86\pm 0.05$$

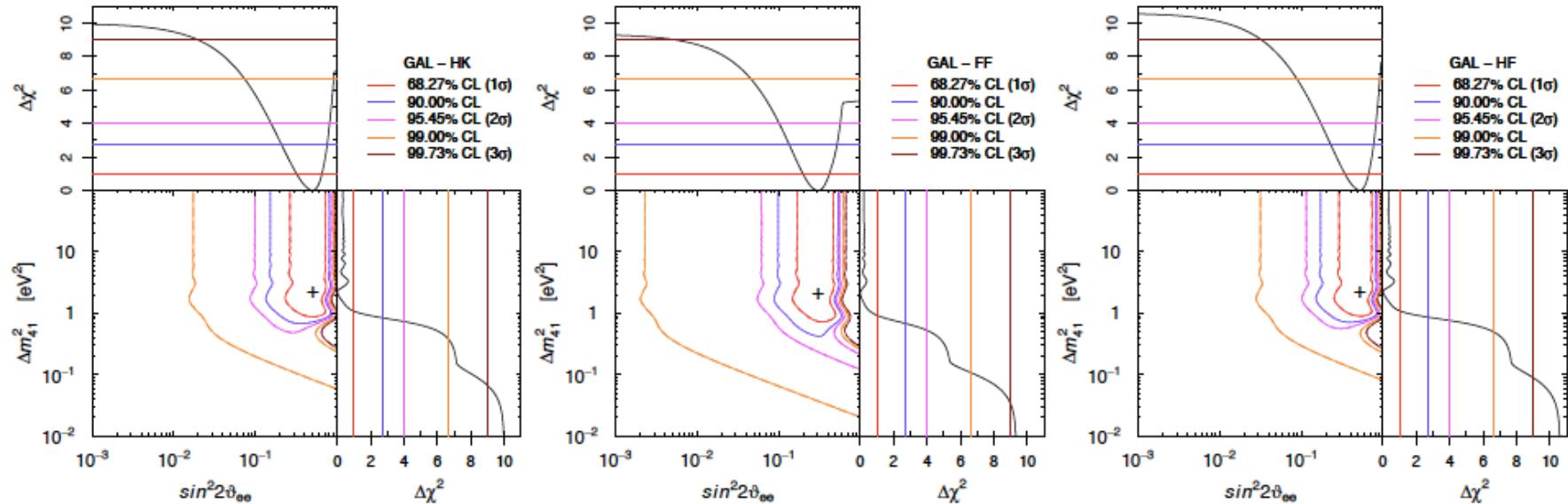
GALLEX & SAGE observe fewer events than expected from their calibration measurements, consistent with ν_e disappearance to sterile neutrinos

Radioactive Neutrino Source Anomaly

Giunti et al.; arXiv:1210.5715

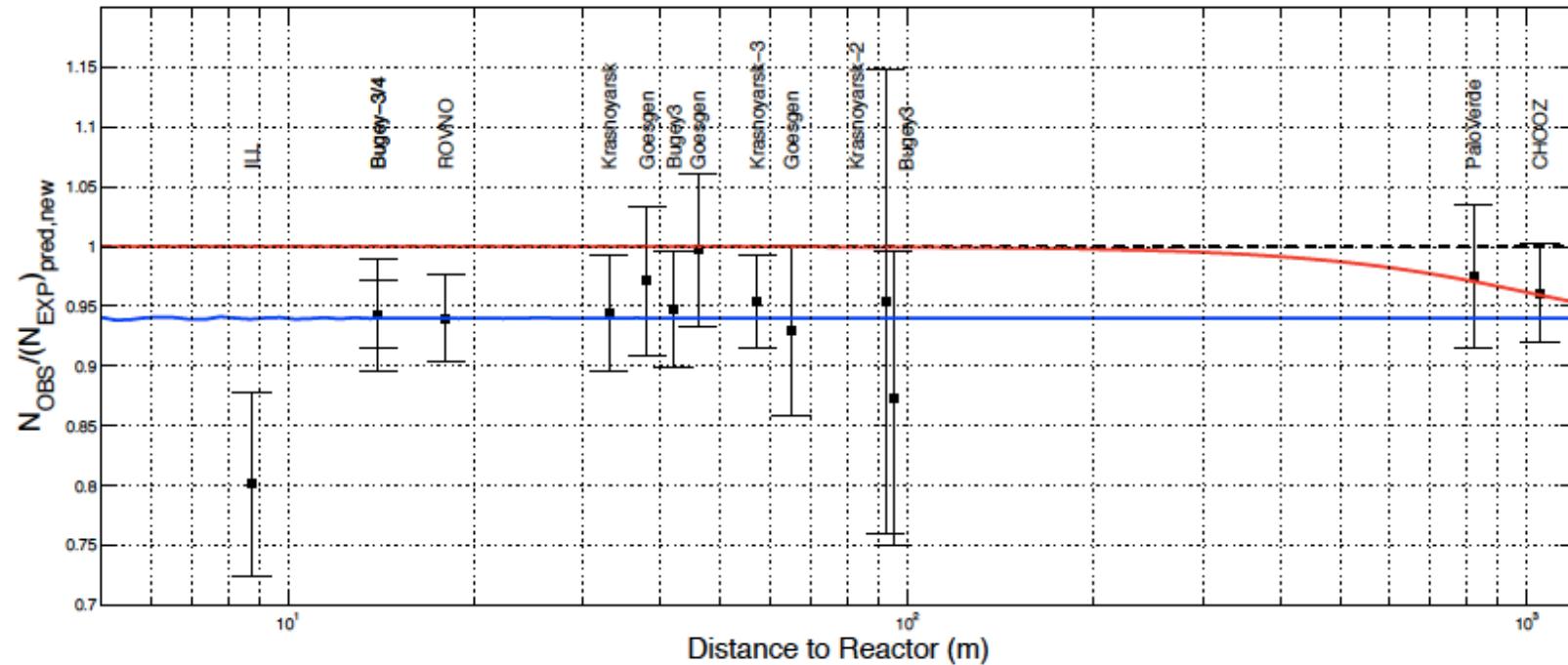
TABLE II. Ratios of measured and expected ^{71}Ge event rates in the four radioactive source experiments. G1 and G2 denote the two GALLEX experiments with ^{51}Cr sources [30–32], S1 denotes the SAGE experiment with a ^{51}Cr source, and S2 denotes the SAGE experiment with a ^{37}Ar source [33–36]. AVE denotes the weighted average.

	G1	G2	S1	S2	AVE
R_B	$0.95^{+0.11}_{-0.11}$	$0.81^{+0.10}_{-0.11}$	$0.95^{+0.12}_{-0.12}$	$0.79^{+0.08}_{-0.08}$	$0.86^{+0.05}_{-0.05}$
R_{HK}	$0.85^{+0.12}_{-0.12}$	$0.71^{+0.11}_{-0.11}$	$0.84^{+0.13}_{-0.12}$	$0.71^{+0.09}_{-0.09}$	$0.77^{+0.08}_{-0.08}$
R_{FF}	$0.93^{+0.11}_{-0.11}$	$0.79^{+0.10}_{-0.11}$	$0.93^{+0.11}_{-0.12}$	$0.77^{+0.09}_{-0.07}$	$0.84^{+0.05}_{-0.05}$
R_{HF}	$0.83^{+0.13}_{-0.11}$	$0.71^{+0.11}_{-0.11}$	$0.83^{+0.13}_{-0.12}$	$0.69^{+0.10}_{-0.09}$	$0.75^{+0.09}_{-0.07}$



Reactor Neutrino Anomaly

G. Mention et al., Phys.Rev.D83:073006,2011

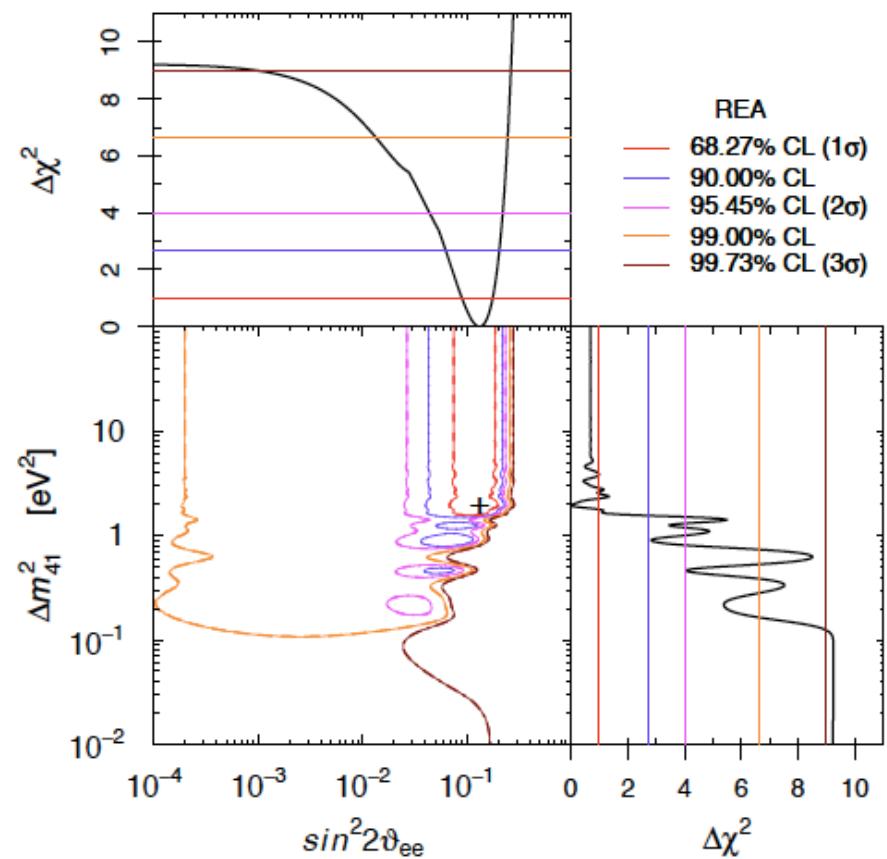
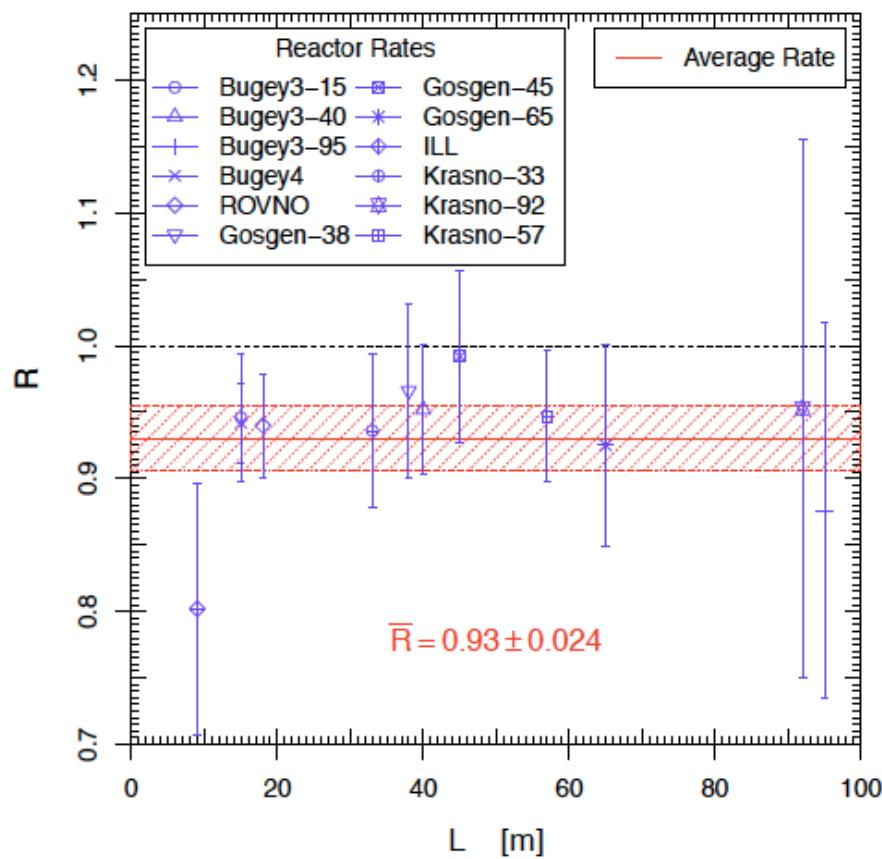


$$R=0.937 \pm 0.027$$

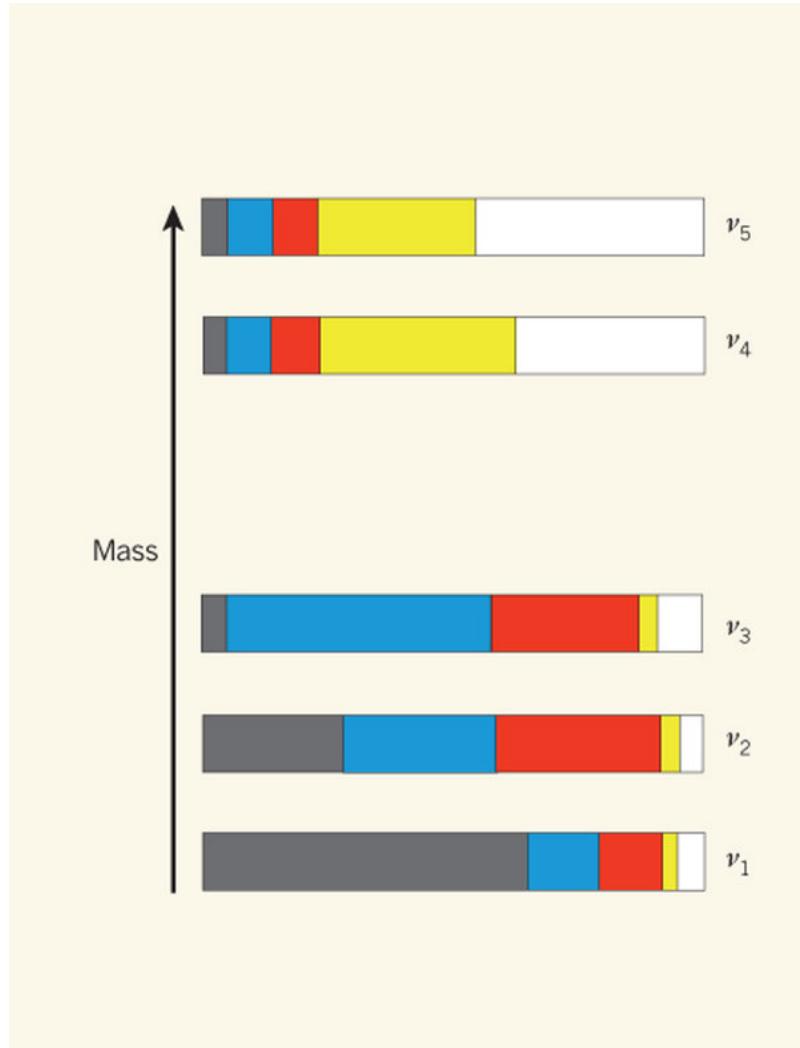
Reactor Neutrino experiments observe fewer events than expected, consistent with $\bar{\nu}_e$ disappearance to sterile neutrinos

Reactor Neutrino Anomaly

Giunti et al.; arXiv:1210.5715



Sterile Neutrinos



- 3+N models
- N>1 allows CP violation for short baseline experiments
 - $\nu_\mu \rightarrow \nu_e \neq \bar{\nu}_\mu \rightarrow \bar{\nu}_e$

3+N Models Requires Large ν Disappearance!

In general, $P(\nu_\mu \rightarrow \nu_e) < \frac{1}{4} P(\nu_\mu \rightarrow \nu_x) P(\nu_e \rightarrow \nu_x)$

Reactor Experiments: $P(\nu_e \rightarrow \nu_x) \sim 15\%$

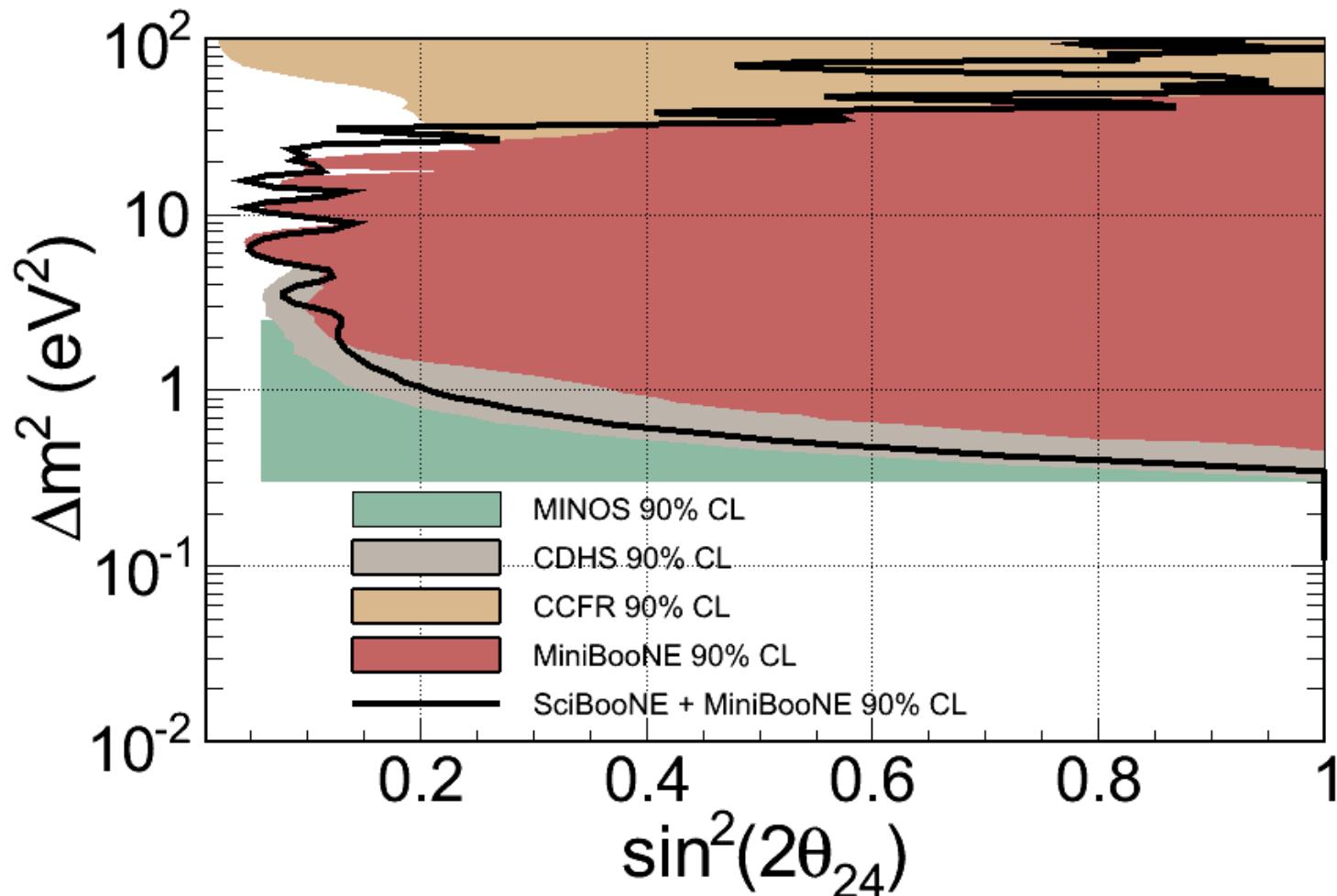
LSND/MiniBooNE: $P(\nu_\mu \rightarrow \nu_e) \sim 0.25\%$

Therefore: $P(\nu_\mu \rightarrow \nu_x) > 7\%$

Assuming that the 3 light neutrinos are mostly active
and the N heavy neutrinos are mostly sterile.

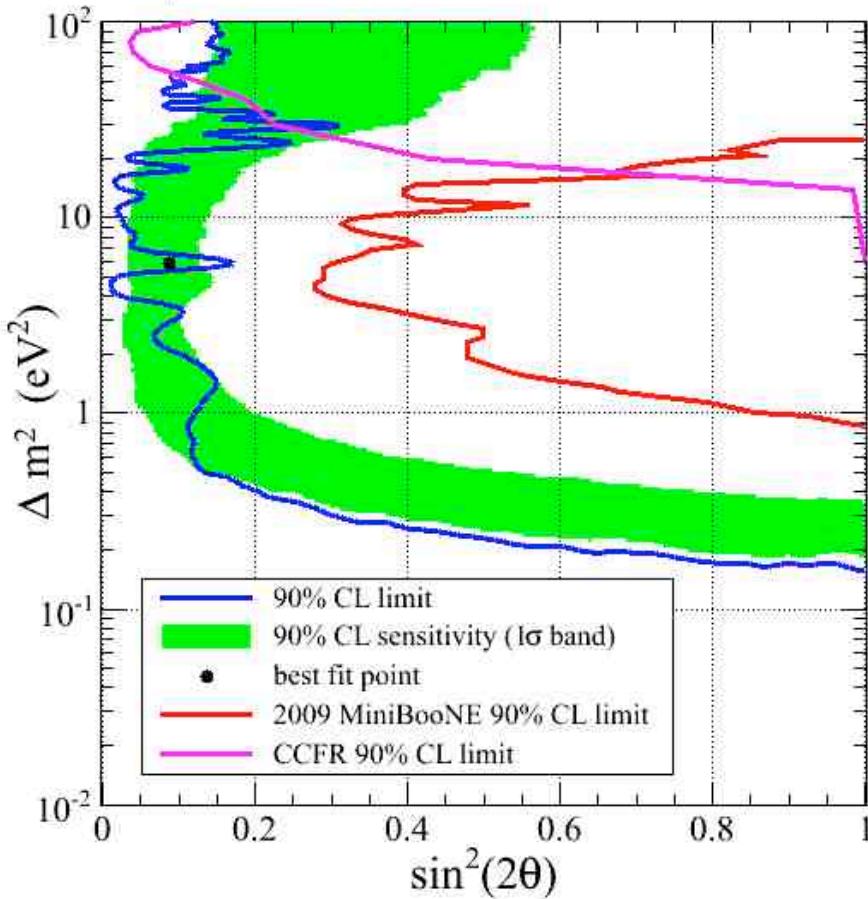
MINOS Limits on ν_μ Disappearance

Alexandre Sousa, arXiv:1110.3455



ν_μ Disappearance Results

Joint $\bar{\nu}_\mu$ Disappearance Analysis



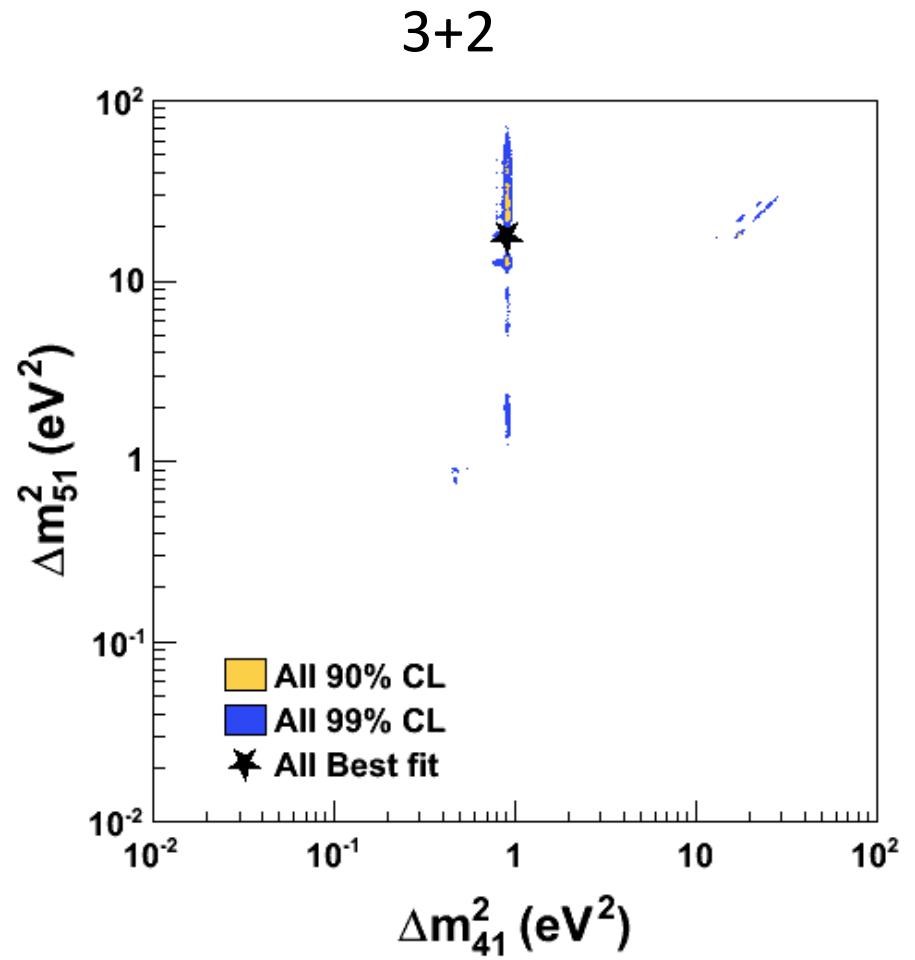
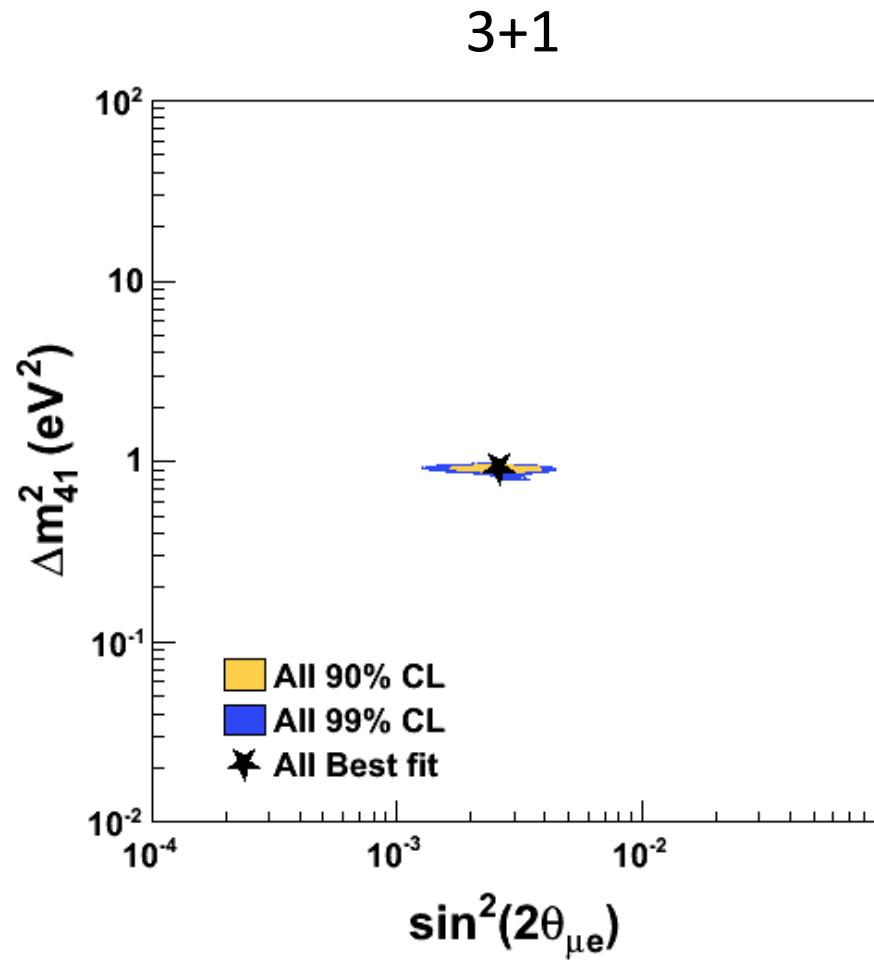
MiniBooNE + SciBooNE

- Consistent with no disappearance
- Best fit point: $\Delta m^2 = 5.9$ eV 2 , $\sin^2 2\theta = 0.086$
- $\chi^2 = 40.0$ (probability 47.1%) at the best fit point
- $\chi^2 = 43.5$ (probability 41.2%) for the null hypothesis
- With $\Delta\chi^2 = 3.5$, null is excluded at 81.9% confidence level
- Probabilities are based on fake data studies

G. Cheng et al., Phys. Rev. D86, 052009 (2012)

Global 3+N Fits to World Data

J.M. Conrad, C.M. Ignarra, G. Karagiorgi, M.H. Shaevitz, & J. Spitz, arXiv:1207.4765



Conclusion

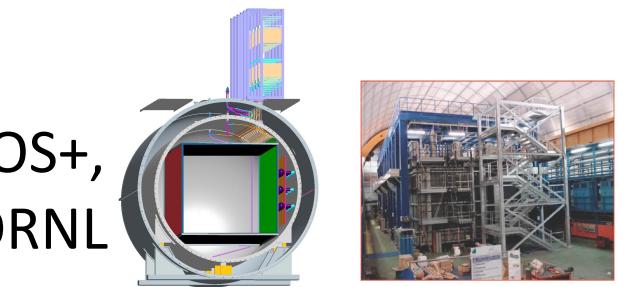
- The anomalies in short baseline ν experiments cannot be explained by the 3 ν paradigm and suggest the existence of sterile ν .
- Sterile ν would contribute to the dark matter of the universe and would have a big impact on astrophysics and cosmology.
- The world neutrino & antineutrino data can be fit fairly well to a 3+N oscillation model, although there is some tension between appearance and disappearance experiments.
- Upcoming experiments (over a wide range of energies) have the potential of proving whether light, sterile neutrinos exist!

Backup

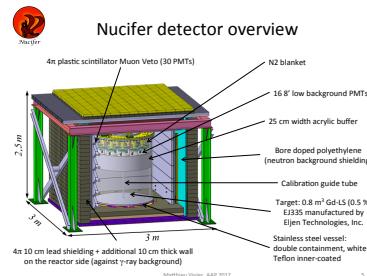
Future ν Experiments

- There is a diverse set of experiments, spanning vastly different energy Scales (from ~ 1 MeV to ~ 10 TeV), that have been proposed to test the 3+N models & resolve the present anomalies:

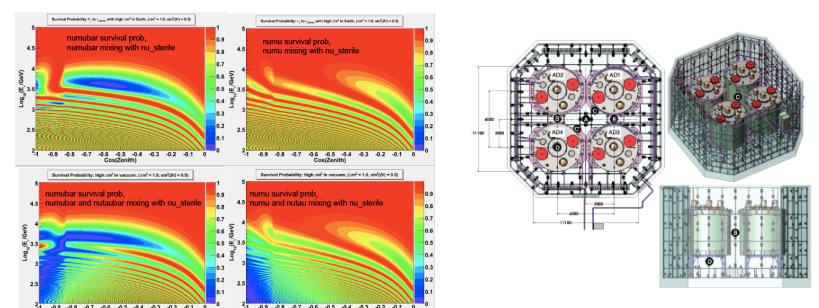
- Accelerator ν Experiments: MicroBooNE, MINOS+, NuStorm at FNAL, ICARUS at CERN, OscSNS at ORNL or J-PARC, IsoDAR



- Reactor ν Experiments:
SCRAAM, NUCIFER, Stereo



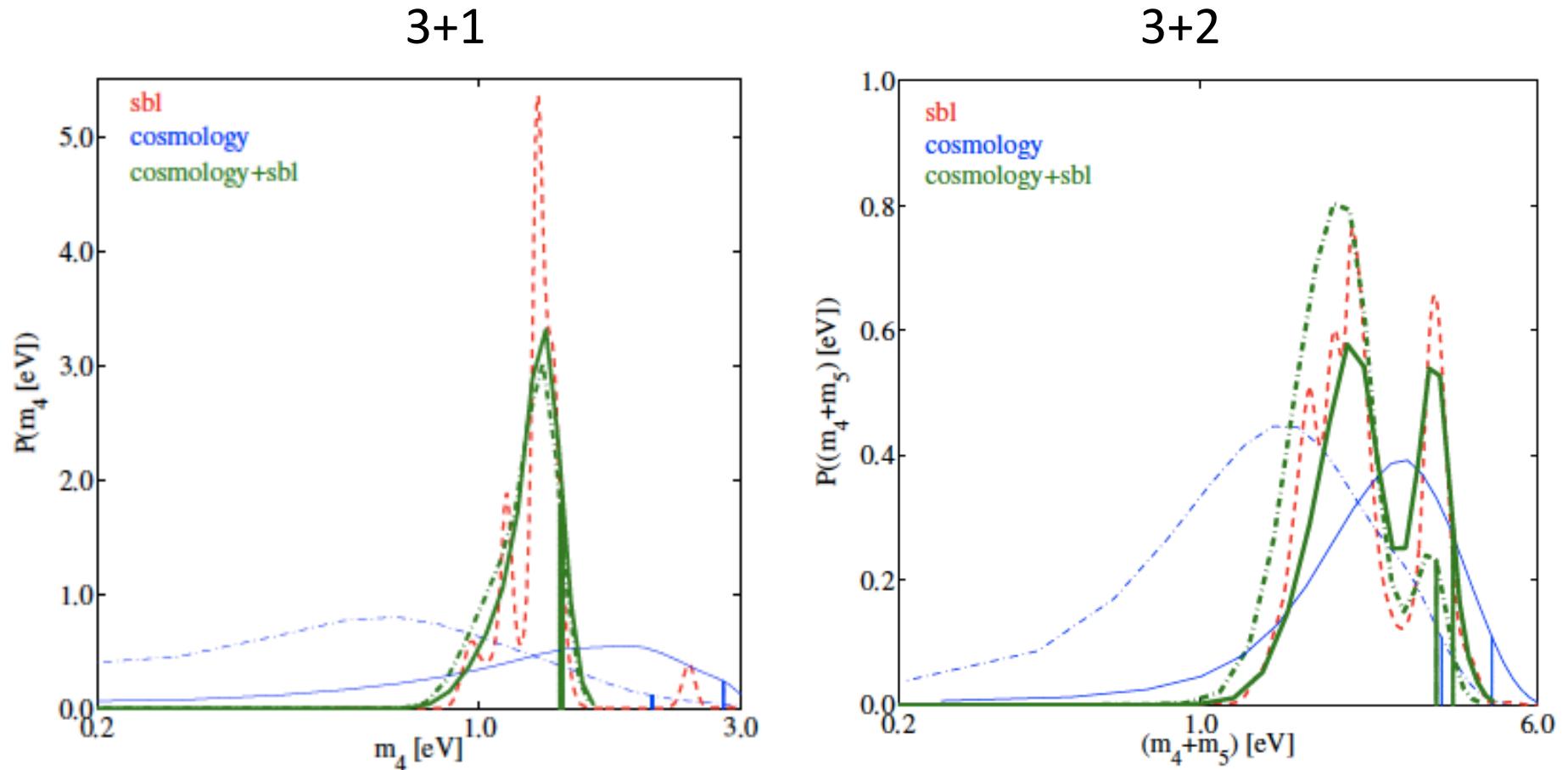
- Radioactive Source ν Experiments: BOREXINO, KamLAND, Daya Bay, Baksan, LENS



- Atmospheric ν Experiments: IceCube

Global 3+N Fits to World Data

Maria Archidiacono, Nicolao Fornengo, Carlo Giunti, Steen Hannestad, & Alessandro Melchiorri
arXiv:1302.6720



Global 3+N Fits to World Data

J.M. Conrad, C.M. Ignarra, G. Karagiorgi, M.H. Shaevitz, & J. Spitz, arXiv:1207.4765

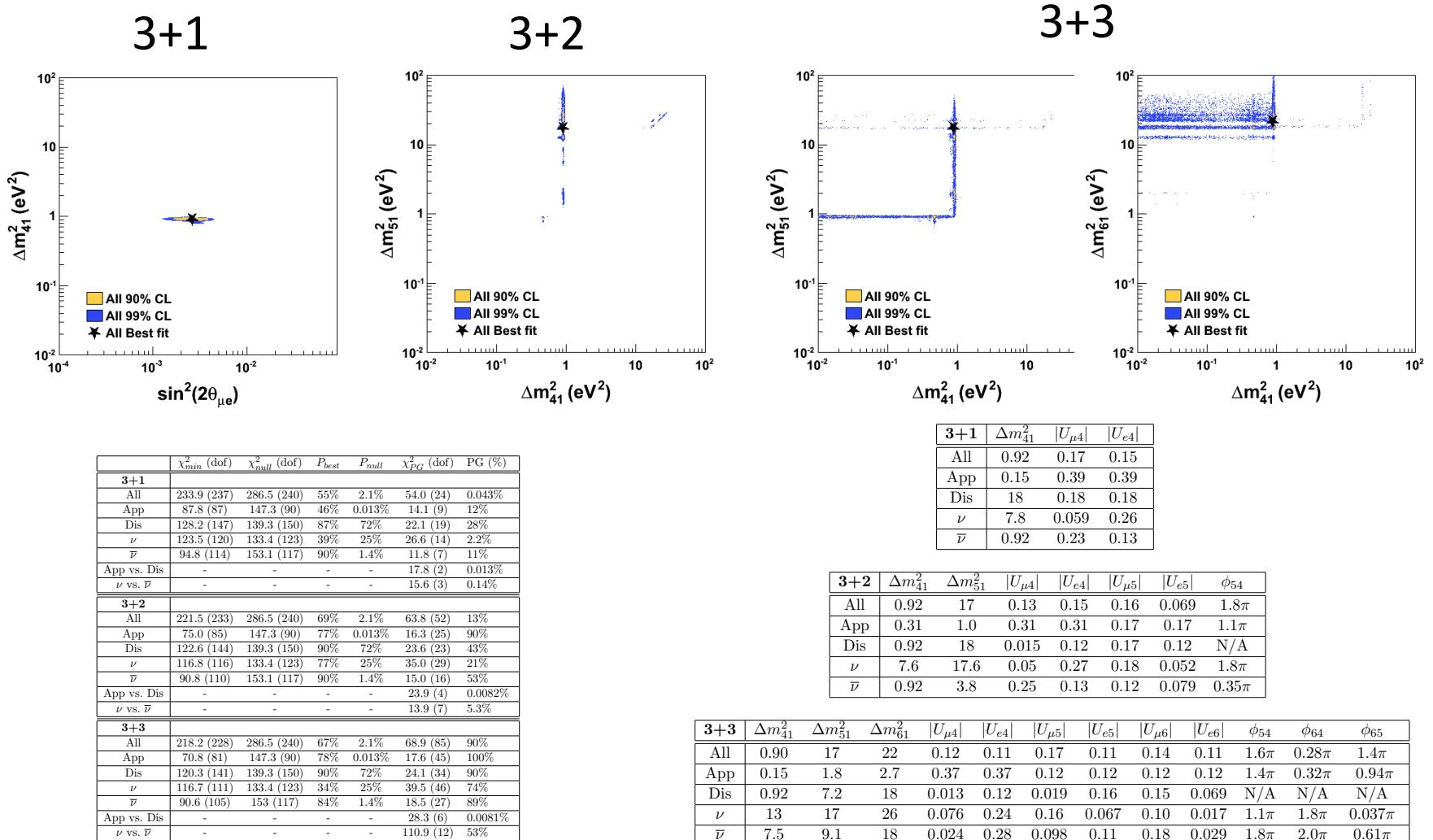
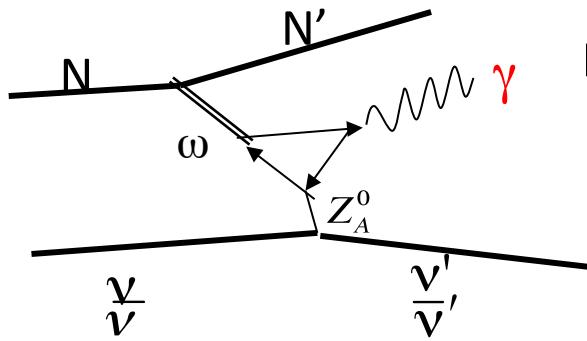


Table 2: The χ^2 values, degrees of freedom (dof) and probabilities associated with the best-fit and null hypothesis in each scenario. Also shown are the results from the Parameter Goodness-of-fit tests. P_{best} refers to the χ^2 -probability at the best fit point and P_{null} refers to the χ^2 -probability at null.

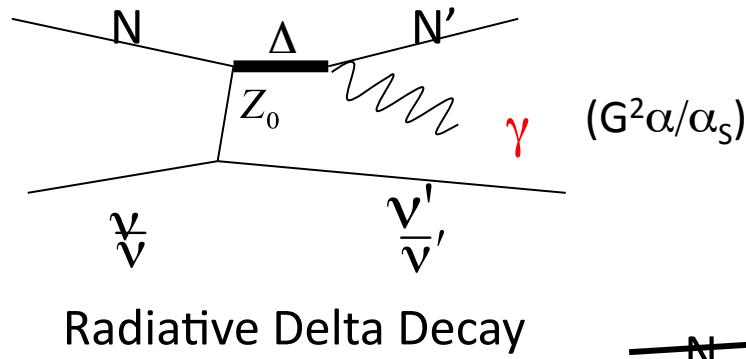
Table 3: The oscillation parameter best-fit points in each scenario considered. The values of Δm^2 shown are in units of eV²

NC γ Backgrounds: Order $(G^2\alpha\alpha_s)$, single γ FS?

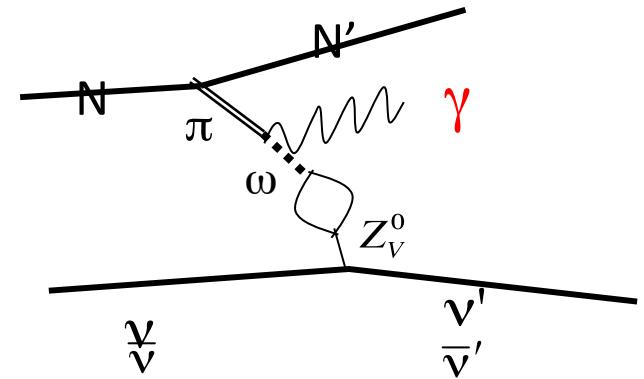
Dominant process
accounted for in MC!



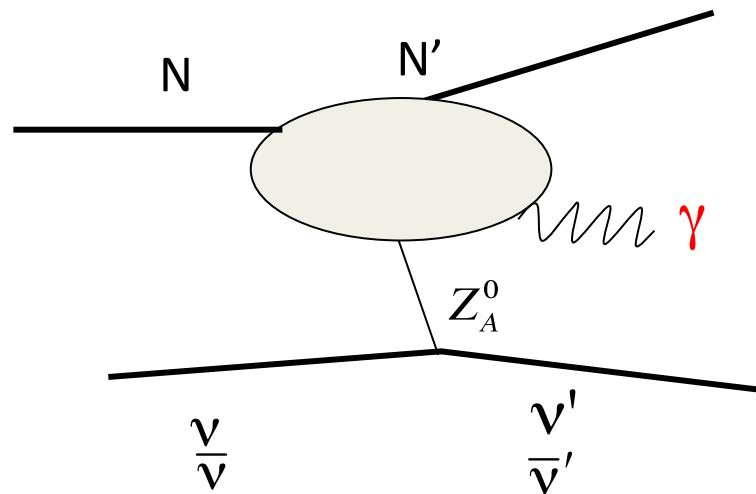
Axial Anomaly



Radiative Delta Decay



Other PCAC



So far no one has found a NC process to account for the ν low-energy excess. Work is in progress:
R. Hill, arXiv:0905.0291
Jenkins & Goldman, arXiv:0906.0984
Zhang & Serot, arXiv:1210.3610

[arXiv:1210.3610](http://arxiv.org/abs/1210.3610)

Multi-Nucleon Nuclear Effects & Neutrino Energy

Martini, Ericson, & Chanfray, arXiv:1211.1523

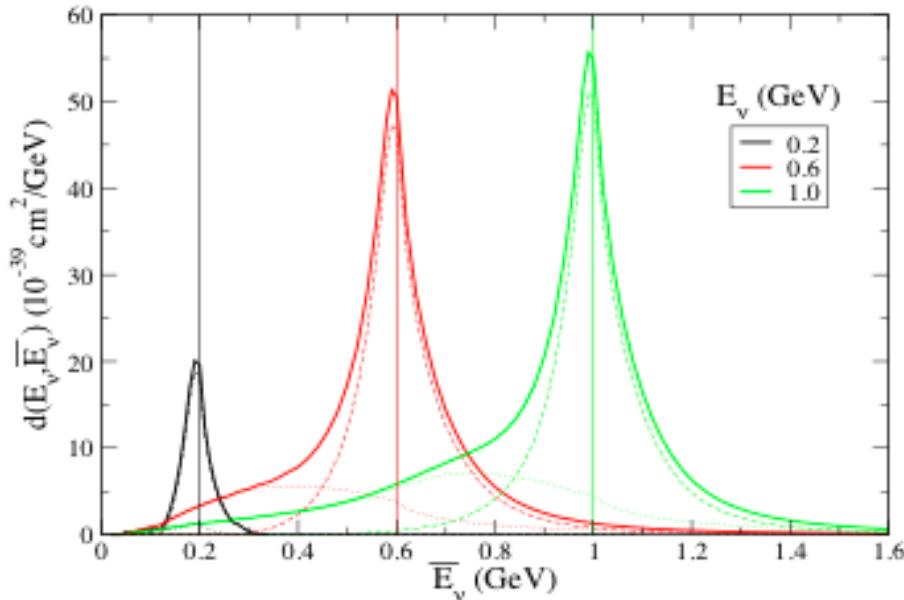


FIG. 1: (Color online) The spreading function $d(E_\nu, \bar{E}_\nu)$ of Eq. (1) per neutron of ^{12}C in the case of electrons evaluated for three E_ν values. The genuine quasielastic (dashed lines) and the multinucleon (dotted lines) contributions are also shown separately.

Multi-Nucleon Nuclear Effects cause MiniBooNE to underestimate E_ν !

Nominal best antineutrino fit: $\Delta m^2 = 0.043 \text{ eV}^2$, $\sin^2 2\theta = 0.88$

Martini Inspired Model: $\Delta m^2 = 0.059 \text{ eV}^2$, $\sin^2 2\theta = 0.64$

ν_μ & ν_e Disappearance

- MB fits performed to date assume small ν_e & ν_μ disappearance
- However, ν_e (ν_μ) disappearance in 3+N models will cause the intrinsic ν_e background to be overestimated (underestimated)
- Therefore, MB is now working on fitting Δm^2 & both U_{e4} & $U_{\mu 4}$:

$$\sin^2 2\theta_{\mu e} = 4(U_{e4} U_{\mu 4})^2$$

$$\sin^2 2\theta_{\mu \mu} = 4(U_{\mu 4})^2 (1 - (U_{\mu 4})^2)$$

$$\sin^2 2\theta_{\mu e} = 4(U_{e4})^2 (1 - (U_{e4})^2)$$

Nominal best antineutrino fit: $\Delta m^2 = 0.043 \text{ eV}^2$, $\sin^2 2\theta = 0.88$

Model with Disappearance: $\Delta m^2 = 0.177 \text{ eV}^2$, $\sin^2 2\theta = 0.07$